

11.6 & 15) THE POWER THAT SHOULD BE RADIATED BY THE e^- IN THE BOHR ORBIT, $P = \frac{2kq^2 a^2}{3c^3}$

SHOULD CAUSE THE ATOM TO COLLAPSE. FIND THE TIME.

a) USING THE CHAIN RULE

$$\frac{dr}{dt} = \frac{dr}{dE} \frac{dE}{dt}$$

FOR A CIRCULAR ORBIT, $E = -\frac{1}{2} \frac{ke^2}{r}$ AND THE RADIATED POWER IS ENERGY LOST BY THE e^- , SO

$$E = -\frac{1}{2} \frac{ke^2}{r} \Rightarrow r = \frac{-ke^2}{2E} \quad \frac{dr}{dE} = + \frac{ke^2}{2E^2}$$

SUBSTITUTING E INTO THE RESULT,

$$\frac{dr}{dE} = + \frac{ke^2}{2} \left(\frac{2r}{ke^2} \right)^2 = \frac{2r^2}{ke^2}$$

THEN

$$\frac{dE}{dt} = - \frac{2ke^2 a^2}{3c^3}$$

FOR THE e^- IN ITS ORBIT

$$\frac{ke^2}{r^2} = m_e a_c \Rightarrow a_c = \frac{ke^2}{m_e r^2}$$

GIVING

$$\frac{dE}{dt} = - \frac{2ke^2}{3c^3} \left(\frac{ke^2}{m_e r^2} \right)^2 \frac{c^2}{c^2}$$

TO CHANGE FROM m_e TO $m_e c^2$

$$\frac{dE}{dt} = - \frac{2(ke^2)^3 c}{3(m_e c^2)^2 r^4}$$

Thus

$$\frac{dr}{dt} = \left(\frac{2r^2}{ke^2} \right) \left(- \frac{2(ke^2)^3 c}{3(m_e c^2)^2 r^4} \right) = - \frac{4(ke^2)^2 c}{3(m_e c^2)^2 r^2}$$

For $r = a_0$

$$\frac{dr}{dt} = - \frac{4(ke^2)^2 c}{3(m_e c^2)^2 a_0^2}$$

THE RATE AT WHICH THE e^- FALLS IN TO THE PROTON IN H.

11.6 & 15) CONTINUED

b) ASSUME $\frac{dr}{dt} = \text{CONSTANT} = -A$ FIND t TO COLLAPSE

$$\frac{dr}{dt} = -A \Rightarrow dt = -\frac{1}{A} dr$$

$$t_{\text{collapse}} = -\frac{1}{A} \int_{a_0}^0 dr = -\frac{r}{A} \Big|_{a_0}^0 = \frac{a_0}{A}$$

EVALUATING

$$a_0 = 5.29 \times 10^{-11} \text{ m} \left(\frac{\text{fm}}{1 \times 10^{-15} \text{ m}} \right)$$

$$a_0 = 5.29 \times 10^4 \text{ fm}$$

$$t_{\text{collapse}} = a_0 \left(\frac{3(m_e c^2)^2 a_0^2}{4(k_e^2)^2 c} \right)$$

$$= \frac{3(m_e c^2)^2 a_0^3}{4(k_e^2)^2 c} = \frac{3(0.511 \text{ MeV})^2 (5.29 \times 10^4 \text{ fm})^2}{4(1.44 \text{ MeV} \cdot \text{fm})^2 (3 \times 10^8 \text{ m/s}^2)} (5.29 \times 10^{-11} \text{ m})$$

$$\boxed{t_{\text{collapse}} = 4.63 \times 10^{-11} \text{ s}} \text{ QUICK! So WTF IS THIS SILVER?}$$

c) MAKE A BETTER APPROXIMATION BY RECOGNIZING

$$\frac{dr}{dt} = \frac{dr}{dt} \Rightarrow dt = \frac{dr}{dr/dt} \Rightarrow t_{\text{collapse}} = \int_{a_0}^0 \frac{dr}{dr/dt}$$

RE-WRITE dr/dt WITH r INSTEAD OF a_0

$$t_{\text{collapse}} = -\frac{3(m_e c^2)^2}{4(k_e^2)^2 c} \int_{a_0}^0 r^2 dr = -\frac{3(m_e c^2)^2}{4(k_e^2)^2 c} \left[\frac{r^3}{3} \Big|_{a_0}^0 \right]$$

$$t_{\text{collapse}} = \frac{3(m_e c^2)^2}{4(k_e^2)^2 c} \frac{a_0^3}{3} = \frac{(m_e c^2)^2}{4(k_e^2)^2 c} a_0^3$$

SO THIS IS ONE THIRD OF THE PREVIOUS RESULT

$$\boxed{t_{\text{collapse}} = 1.54 \times 10^{-11} \text{ s}} \text{ STILL QUICK!}$$