

- 2) a) FIND THE POPULATIONS OF THE 1ST AND 2ND EXCITED STATES RELATIVE TO THE GROUND STATE FOR ATOMIC H AT 300K, ASSUMING M-B STATISTICS ARE VALID.
- b) FIND THE SAME POPULATIONS FOR η UMa AT $T = 15,500$ K.
- c) FOR THE POPULATIONS IN b, DETERMINE IF BALMER OR PASCHEN ABSORPTION LINES WILL BE STRONGER.

a) THE RATIO OF POPULATIONS IS

$$\frac{n(E_3)}{n(E_1)} = \frac{g(E_3) \lambda e^{-\beta E_3}}{g(E_1) \lambda e^{-\beta E_1}} = \frac{g(E_3)}{g(E_1)} e^{(E_1 - E_3)/kT}$$

FOR HYDROGEN

$$n=1: \quad g(E_1) = 2 \quad E_1 = -13.6 \text{ eV} \quad (1s)$$

$$n=2 \quad g(E_2) = 8 \quad E_2 = -3.4 \text{ eV} \quad (2s + 2p)$$

$$n=3 \quad g(E_3) = 18 \quad E_3 = -1.51 \text{ eV} \quad (3s + 3p + 3d)$$

THUS

$$\frac{n_3}{n_1} = \frac{18}{2} e^{(-13.6 + 1.5)/(8.62 \times 10^{-5})(300)} = 9 e^{-468} \approx 0!$$

$$\frac{n_2}{n_1} = \frac{8}{2} e^{(-13.6 + 3.4)/(8.62 \times 10^{-5})(300)} = 4 e^{-469} \approx 0!$$

\Rightarrow AT ROOM TEMP., VERY FEW H ATOMS ARE EXCITED!

b) FOR η UMa AT $T = 15,500$ K,

$$\frac{n_3}{n_1} = 9 e^{(-12.1)/(8.62 \times 10^{-5})(15500)} = 9 e^{-9.06} = 0.001$$

$$\frac{n_2}{n_1} = 4 e^{(-10.2)/(8.62 \times 10^{-5})(15500)} = 4 e^{-7.63} = 0.0019$$

c) \Rightarrow SINCE n_2 IS SLIGHTLY MORE POPULATED, TRANSITIONS UPWARD FROM IT WILL BE STRONGER THAN FROM n_3

\Rightarrow BALMER STRONGER THAN PASCHEN!